Hyperonic Three-Body Forces & Consequences for Neutron Stars

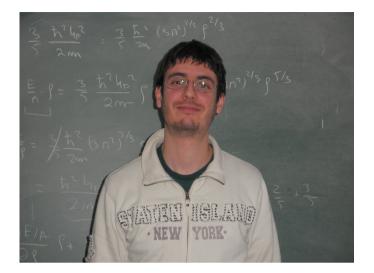
Isaac Vidaña CFisUC, University of Coimbra



Microphysics in Computational Relativistic Astrophysics MICRA 2015 August 17th – 22nd 2015, Stockholm (Sweden) In this talk I will ...

- Present NNA and NNΣ forces based on a two-meson exchange model
- Analyze the role of these forces in the solution of the hyperon puzzle

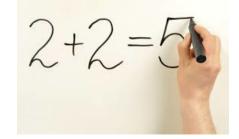
This study was part of the Ph.D. Thesis of Domenico Logoteta (University of Coimbra, September 2013)



Of course ...

- \diamond He deserves all the credits
- ♦ I am the only responsible for any possible mistake





Hyperons in Neutron Stars

Hyperons in NS considered by many authors since the pioneering work of Ambartsumyan & Saakyan (1960)



Phenomenological approaches

- ♦ Relativistic Mean Field Models: Glendenning 1985; Knorren et al. 1995; Shaffner-Bielich & Mishustin 1996, Bonano & Sedrakian 2012, ...
- ♦ Non-realtivistic potential model: Balberg & Gal 1997
- ♦ Quark-meson coupling model: Pal et al. 1999, …
- ♦ Chiral Effective Lagrangians: Hanauske et al., 2000
- ♦ Density dependent hadron field models: Hofmann, Keil & Lenske 2001

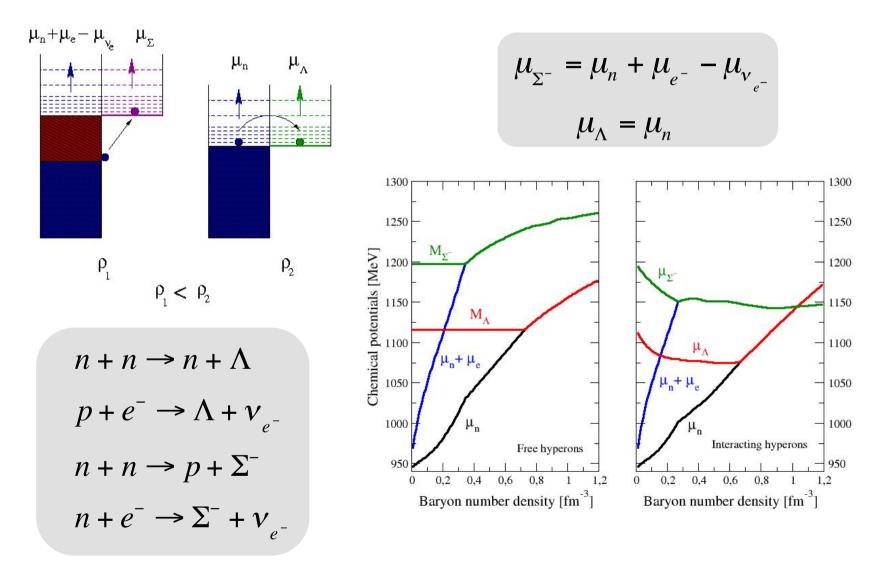


Microscopic approaches

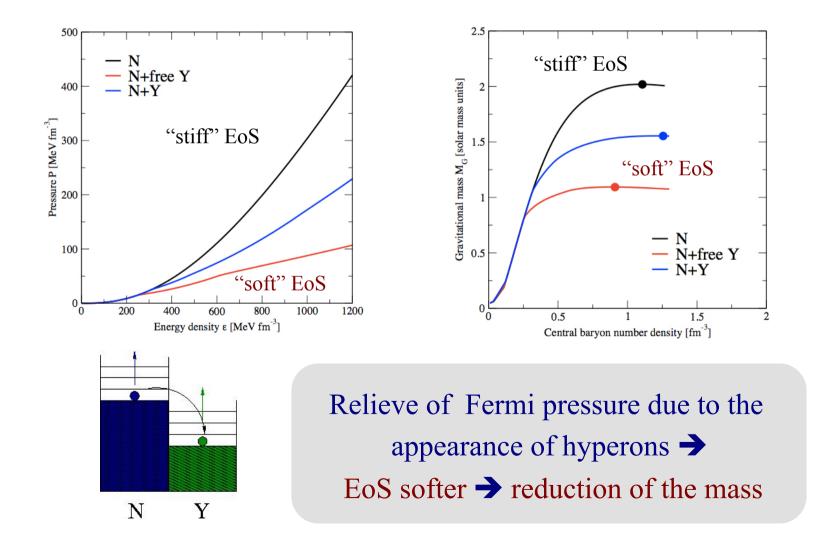
- ♦ Brueckner-Hartree-Fock theory: Baldo et al. 2000; I. V. et al. 2000, Schulze et al. 2006, I.V. et al. 2011, Burgio et al. 2011, Schulze & Rijken 2011
- ♦ DBHF: Sammarruca (2009), Katayama & Saito (2014)
- ♦ V_{low k}: Djapo, Schaefer & Wambach, 2010



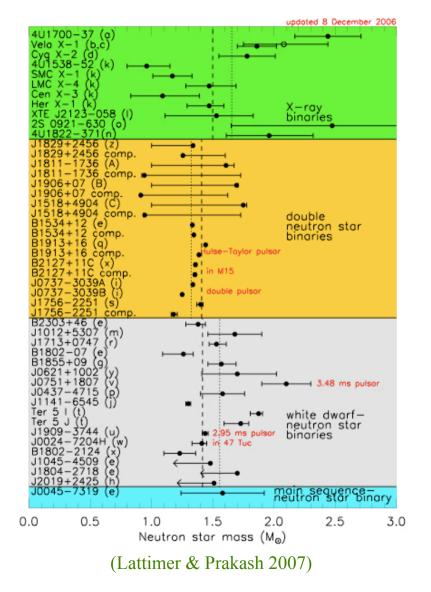
Hyperons are expected to appear in the core of neutron stars at $\rho \sim (2-3)\rho_0$ when μ_N is large enough to make the conversion of N into Y energetically favorable.

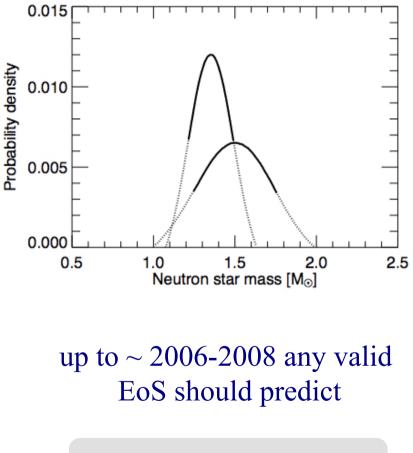


Effect of Hyperons in the EoS and Mass of Neutron Stars



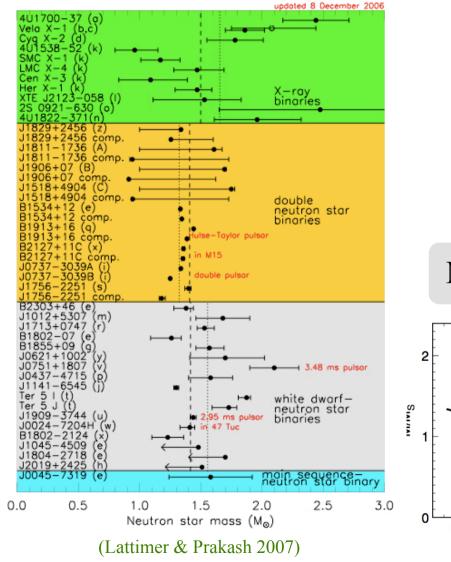
Measured Neutron Star Masses (up to $\sim 2006-2008$)



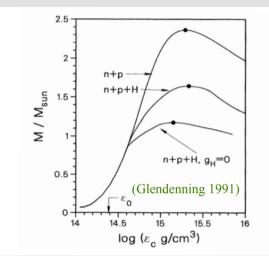


$$M_{\rm max} [EoS] > 1.4 - 1.5 M_{\odot}$$

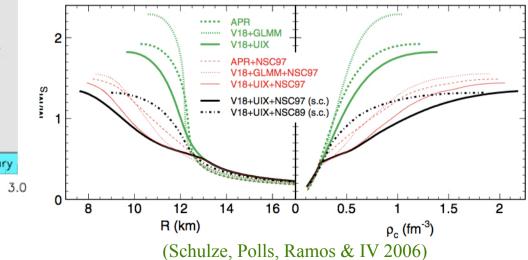
Hyperons in NS (up to ~ 2006-2008)



Phenomenological: M_{max} compatible with 1.4-1.5 M_{\odot}

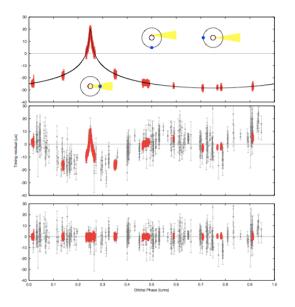


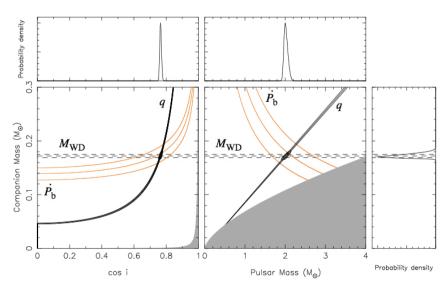
Microscopic : $M_{max} < 1.4-1.5 M_{\odot}$



Recent measurements of high masses —> life of hyperons more difficult

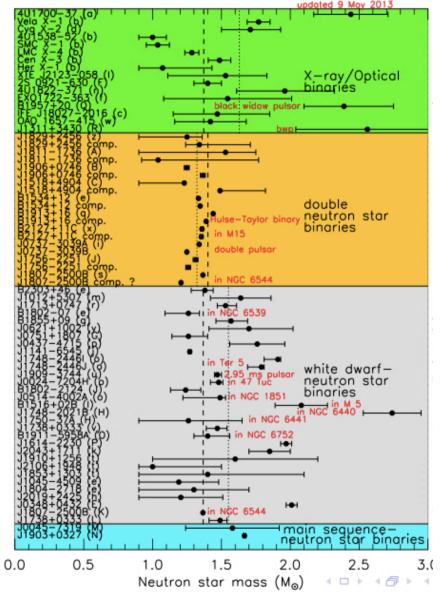
- PSR J164-2230 (Demorest et al. 2010)
 - ✓ binary system (P=8.68 d)
 - ✓ low eccentricity (ϵ =1.3 x 10⁻⁶)
 - \checkmark companion mass: $\sim 0.5 M_{\odot}$
 - ✓ pulsar mass: $M = 1.97 \pm 0.04 M_{\odot}$





- <u>PSR J0348+0432</u> (Antoniadis et al. 2013)
 - ✓ binary system (P=2.46 h)
 - \checkmark very low eccentricity
 - \checkmark companion mass: $0.172 \pm 0.003 M_{\odot}$
 - ✓ pulsar mass: $M = 2.01 \pm 0.04 M_{\odot}$

Measured Neutron Star Masses (2015)



updated from Lattimer 2013

Observation of $\sim 2 M_{sun}$ neutron stars

Dense matter EoS stiff enough is required such that

 $M_{\rm max} [EoS] > 2M_{\odot}$

Can hyperons still be present in the interior of neutron stars in view of this constraint ?

The Hyperon Puzzle



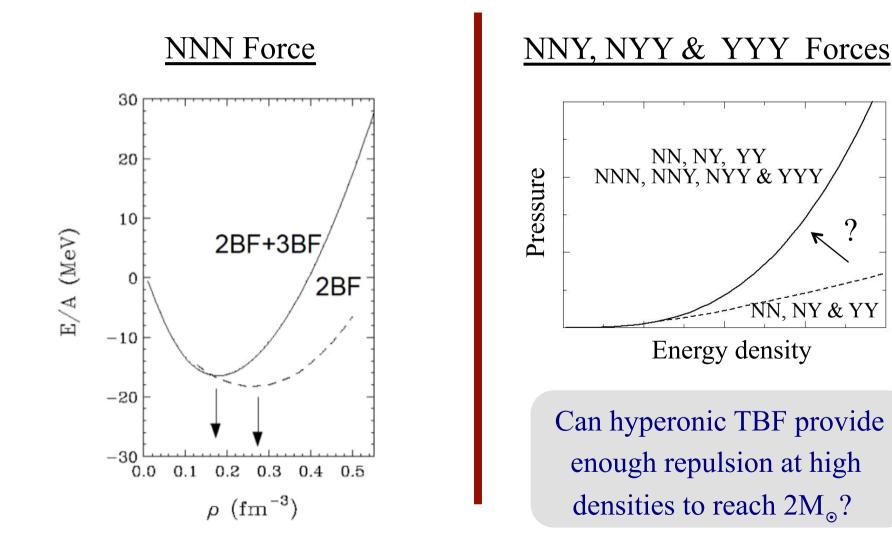
"Hyperons \rightarrow "soft (or too soft) EoS" not compatible (mainly in microscopic approaches) with measured (high) masses. However, the presence of hyperons in the NS interior seems to be unavoidable."



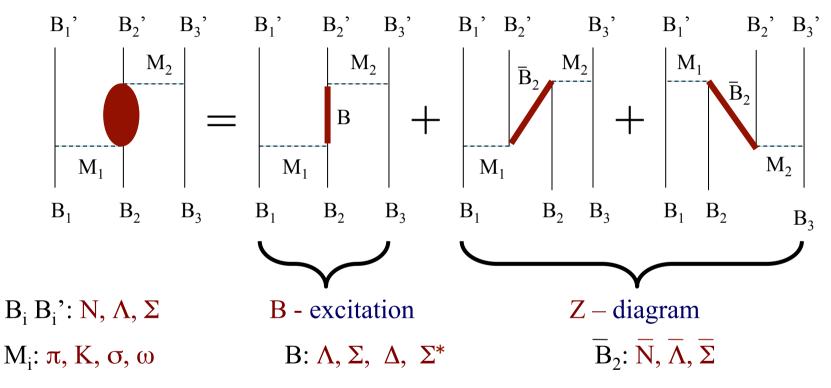
- $\checkmark\,$ can YN & YY interactions still solve it ?
- \checkmark or perhaps hyperonic three-body forces ?
- ✓ what about quark matter ?

Can Hyperonic TBF solve this puzzle?

Natural solution based on: Importance of NNN force in Nuclear Physics (Considered by several authors: Chalk, Gal, Usmani, Bodmer, Takatsuka, Loiseau, Nogami, Bahaduri, IV)



Two-meson exchange Hyperonic TBF



Vertices: consistent with YN and YY

Repulsion at high densities due to Z-diagram as in NNN

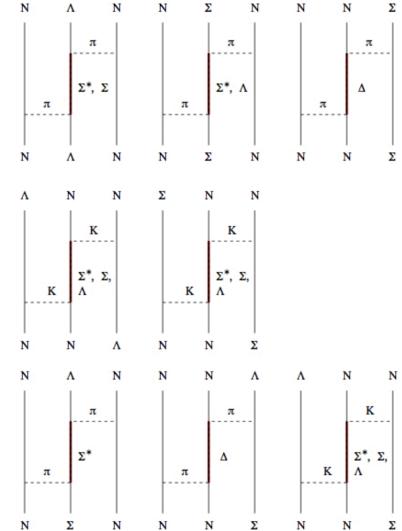
Baryon-excitation contribution

 $(\pi$ -, *K*-exchange)

$$V_{NNY \leftrightarrow NNY'}^{M_1M_2,R} = C_{NNY \leftrightarrow NNY'}^{M_1M_2,R} \left(\hat{O}_I \left\{ X_{12}(\vec{r}_{12}), X_{23}(\vec{r}_{23}) \right\} + \hat{O}_{II} \left[X_{12}(\vec{r}_{12}), X_{23}(\vec{r}_{23}) \right] \right)$$

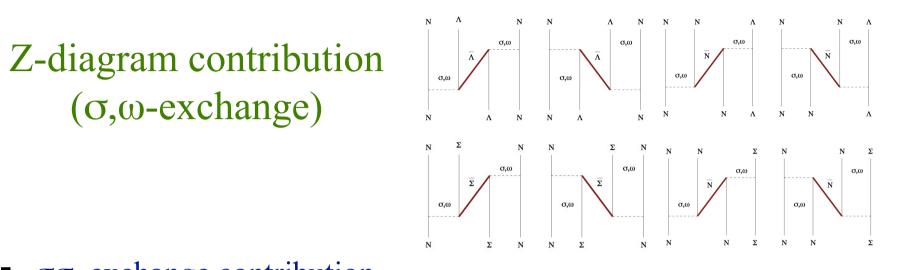
 $\hat{O}_{I}, \hat{O}_{II} \rightarrow \text{isospin structure}$

$$\begin{split} X_{ij}(\vec{x}) &= \vec{\sigma}_i \cdot \vec{\sigma}_j Y_{ij}(x) + \hat{S}_{ij}(\hat{x}) T_{ij}(x) \\ Y_{ij}(x) &= \frac{\partial^2 Z_{ij}}{\partial x^2} + \frac{2}{x} \frac{\partial Z_{ij}}{\partial x}, \ T_{ij}(x) = \frac{\partial^2 Z_{ij}}{\partial x^2} - \frac{1}{x} \frac{\partial Z_{ij}}{\partial x} \\ Z_{12}(x) &= \frac{4\pi}{m_{M_1}} \int \frac{d\vec{k}}{(2\pi)^3} \frac{e^{-i\vec{k}\cdot\vec{x}}}{k^2 + m_{M_1}^2} F_{B_1B_1'M_1}(k^2) F_{B_2BM_1}(k^2) \\ Z_{23}(x) &= \frac{4\pi}{m_{M_2}} \int \frac{d\vec{q}}{(2\pi)^3} \frac{e^{-i\vec{k}\cdot\vec{x}}}{q^2 + m_{M_2}^2} F_{B_3B_3'M_2}(q^2) F_{B_2'BM_2}(q^2) \end{split}$$



Numerical factors & Isospin structure

Process	M_1M_2	R	$C^{M_1M_2,R}_{_{NNY\leftrightarrow NNY'}}$	\hat{O}_{I}	\hat{O}_{II}
$NN\Lambda \leftrightarrow NN\Lambda$	ππ	Σ	$\frac{1}{9} \frac{\tilde{f}_{NN\pi}^2 \tilde{f}_{\Lambda\Sigma\pi}^2}{16\pi^2} \frac{m_{\pi}^2}{m_{\Sigma} - m_{\Lambda}}$	$ec{ au}_1\cdotec{ au}_3$	-
	ππ	Σ*	$-\frac{2}{27}\frac{f_{NN\pi}^2 f_{\Lambda\Sigma^*\pi}^2}{\frac{16\pi^2}{\pi^2}}\frac{m_{\pi}^2}{m_{\pi^*}^2}$	$ec{ au}_1\cdotec{ au}_3$	r
	KK	۸	$-\frac{1}{9}\frac{\tilde{f}_{N\Lambda K}^4}{16\pi^2}\frac{m_K^2}{m_{N-m_{\Lambda}}}\frac{m_K^4}{m_K^4}$	1	-
	KK	Σ	$-\frac{1}{18}\frac{\tilde{f}_{N\Lambda K}^{2}\tilde{f}_{N\Sigma K}^{2}}{16\pi^{2}}\frac{m_{K}^{2}}{m_{N}-m_{\Sigma}}\frac{m_{K}^{4}}{m_{\pi}^{4}}$	$\left\{ \vec{I}_1 \cdot \vec{\tau}_2, \vec{\tau}_2 \cdot \vec{I}_3 \right\}$	$\left[ec{I}_1\cdotec{ au}_2,ec{ au}_2\cdotec{I}_3 ight]$
	KK	Σ*	$-\frac{1}{27}\frac{f_{N\Lambda K}^{2}f_{N\Sigma^{*}K}^{2}}{16\pi^{2}}\frac{m_{K}^{2}}{m_{\Sigma^{*}}-m_{N}}\frac{m_{K}^{4}}{m_{\pi}^{4}}$	$\left\{ \vec{I}_1 \cdot \vec{\tau}_2, \vec{\tau}_2 \cdot \vec{I}_3 \right\}$	$-\frac{1}{2}\left[\vec{I}_{1}\cdot\vec{\tau}_{2},\vec{\tau}_{2}\cdot\vec{I}_{3}\right]$
$NN\Sigma \leftrightarrow NN\Sigma$	ππ	٨	$-\frac{1}{9}\frac{\tilde{f}_{NN\pi}^2\tilde{f}_{\Lambda\Sigma\pi}^2}{16\pi^2}\frac{m_{\pi}^2}{m_{\Sigma}-m_{\Lambda}}$	$ec{ au}_1\cdotec{ au}_3$, — ·
	ππ	Δ	$-\frac{2}{21}\frac{f_{N\Sigma\pi}^2f_{N\Delta\pi}^2}{f_{N\Sigma\pi}^2}\frac{m_{\pi}^2}{m_{\pi}^2}$	$\left\{ ec{ au}_1 \cdot ec{ au}_2, ec{ au}_2 \cdot ec{I}_3 ight\}$	$\frac{1}{4}\left[ec{ au}_1\cdotec{ au}_2,ec{ au}_2\cdotec{I}_3 ight]$
	ππ	Σ*	$-\frac{2}{27} \frac{f_{NN\pi}^2 f_{\Sigma\Sigma^*\pi}^2}{16\pi^2} \frac{m_{\Delta}^2 - m_N}{m_{\Sigma^*}^2 - m_{\Sigma}} \\ -\frac{1}{9} \frac{f_{N\Lambda K}^2 f_{N\Sigma K}^2}{16\pi^2} \frac{m_K^2}{m_N - m_{\Lambda}} \frac{m_K^4}{m_{\pi}^4}$	$\vec{ au}_1 \cdot \vec{ au}_3$	
	KK	۸	$-\frac{1}{9} \frac{\tilde{f}_{N\Lambda K}^2 \tilde{f}_{N\Sigma K}^2}{16\pi^2} \frac{m_K^2}{m_N - m_\Lambda} \frac{m_K^4}{m_A^4}$	$ec{ au}_1 \cdot ec{ au}_3$	
	KK	Σ	$-\frac{1}{18}\frac{\tilde{f}_{N\Sigma K}^4}{16\pi^2}\frac{m_K^2}{m_N-m_\Sigma}\frac{m_K^4}{m_K^4}$	$\left\{ \vec{\tau}_1 \cdot \vec{\tau}_2, \vec{\tau}_2 \cdot \vec{\tau}_3 \right\}$	$\left[ec{ au}_1 \cdot ec{ au}_2, ec{ au}_2 \cdot ec{ au}_3 ight]$
	KK	Σ*	$-\frac{1}{27}\frac{\tilde{f}_{N\Sigma K}^{2}f_{N\Sigma^{\star}K}^{2}}{16\pi^{2}}\frac{m_{K}^{2}}{m_{\Sigma^{\star}}-m_{N}}\frac{m_{K}^{4}}{m_{\pi}^{4}}$	$\left\{\vec{\tau}_1\cdot\vec{\tau}_2,\vec{\tau}_2\cdot\vec{\tau}_3\right\}$	$-\frac{1}{2}\left[\vec{\tau}_1\cdot\vec{\tau}_2,\vec{\tau}_2\cdot\vec{\tau}_3\right]$
$NN\Sigma \leftrightarrow NN\Lambda$	ππ	Σ*	$-\frac{2}{27}\frac{\tilde{f}_{NN\pi}^{2}f_{\Lambda\Sigma^{*}\pi}f_{\Sigma\Sigma^{*}\pi}}{16\pi^{2}}\frac{m_{\pi}^{2}}{(m_{\Sigma^{*}}-(m_{\Sigma}+m_{\Lambda})/2)}$	$(\vec{\tau}_1\cdot\vec{I}_2)(\vec{ ho}_2\cdot\vec{ au}_3)$	
	ππ	Δ		$\left\{ ec{ au}_1 \cdot ec{ au}_2, ec{ au}_2 \cdot ec{ au}_3 ight\}$	$\frac{1}{4} \left[\vec{\tau}_1 \cdot \vec{\tau}_2, \vec{\tau}_2 \cdot \vec{\rho}_3 \right]$
	KK	۸	$-\frac{1}{9} \frac{f_{N\Lambda K}^3 f_{N\Sigma K}}{16\pi^2} \frac{m_K^2}{m_N - m_\Lambda} \frac{m_K^4}{m_\pi^4}$	$\vec{\rho}_1 \cdot \vec{\tau}_3$	-
	KK	Σ		$\left\{ ec{ ho}_1\cdotec{ au}_2,ec{ au}_2\cdotec{ au}_3 ight\}$	$\left[ec{ ho}_1 \cdot ec{ au}_2, ec{ au}_2 \cdot ec{ au}_3 ight]$
	KK	Σ*	$-\frac{1}{27}\frac{\tilde{f}_{N\Sigma K}^3 \tilde{f}_{N\Lambda K}^2}{16\pi^2}\frac{m_N^2}{m_{\Sigma^*}-m_N}\frac{m_K^4}{m_{\pi}^4}$		$-\frac{1}{2}\left[\vec{\rho}_1\cdot\vec{\tau}_2,\vec{\tau}_2\cdot\vec{\tau}_3\right]$



 $\sigma\sigma$ -exchange contribution

$$\begin{split} V_{NNY}^{\sigma\sigma,\bar{B}} &= C_{NNY}^{\sigma\sigma,\bar{B}} \left(-4Z_{12}(r_{12})Z_{23}(r_{23})\nabla_{r_{2}}^{2} - 4Z_{12}'(r_{12})Z_{23}(r_{23})\hat{r}_{12} \cdot \nabla_{r_{2}'} \right. \\ &- 4Z_{12}(r_{12})Z_{23}'(r_{23})\hat{r}_{23} \cdot \nabla_{r_{2}'} - \left(Y_{12}(r_{12})Z_{23}(r_{23}) + Z_{12}(r_{12})Y_{23}(r_{23})\right) \\ &- \hat{r}_{12} \cdot \hat{r}_{23}Z_{12}'(r_{12})Z_{23}'(r_{23}) - 2i\left(Z_{12}'(r_{12})Z_{23}(r_{23})\vec{\sigma}_{2} \cdot \hat{r}_{12} \times \nabla_{r_{2}'} + Z_{12}(r_{12})Z_{23}'(r_{23})\vec{\sigma}_{2} \cdot \hat{r}_{23} \times \nabla_{r_{2}'}\right) \right) \delta\left(\vec{r}_{1} - \vec{r}_{1}'\right) \delta\left(\vec{r}_{2} - \vec{r}_{2}'\right) \delta\left(\vec{r}_{3} - \vec{r}_{3}'\right) \end{split}$$

• $\omega\omega$ -exchange contribution

$$\begin{aligned} V_{NNY}^{\omega\omega,\bar{B}} &= C_{NNY}^{\omega\omega,\bar{B}} \left(\left(\left(1 + \vec{\sigma}_1 \cdot \vec{\sigma}_2 + \vec{\sigma}_1 \cdot \vec{\sigma}_2 + \vec{\sigma}_2 \cdot \vec{\sigma}_3 \right) \hat{r}_{12} \cdot \hat{r}_{23} - \vec{\sigma}_1 \cdot \hat{r}_{23} \vec{\sigma}_2 \cdot \hat{r}_{12} - \vec{\sigma}_2 \cdot \hat{r}_{23} \vec{\sigma}_3 \cdot \hat{r}_{12} \right) \\ &- \vec{\sigma}_1 \cdot \hat{r}_{23} \vec{\sigma}_3 \cdot \hat{r}_{12} \right) Z_{12}'(r_{12}) Z_{23}'(r_{23}) - 2i Z_{12}'(r_{12}) Z_{23}(r_{23}) \left(\vec{\sigma}_2 + \vec{\sigma}_3 \right) \cdot \hat{r}_{12} \times \nabla_{r_3'} \\ &- 2i Z_{12}(r_{12}) Z_{23}'(r_{23}) \left(\vec{\sigma}_2 + \vec{\sigma}_3 \right) \cdot \hat{r}_{23} \times \nabla_{r_2'} - 4 Z_{12}(r_{12}) Z_{23}(r_{23}) \nabla_{r_1'} \cdot \nabla_{r_3'} \right) \\ &\delta \left(\vec{r}_1 - \vec{r}_1' \right) \delta \left(\vec{r}_2 - \vec{r}_2' \right) \delta \left(\vec{r}_3 - \vec{r}_3' \right) \end{aligned}$$

• $\sigma \omega$ -exchange contribution

$$\begin{split} W_{NNY}^{\sigma\omega,\bar{B}} &= C_{NNY}^{\sigma\omega,\bar{B}} \left(\left(\left(1 + \vec{\sigma}_{2} \cdot \vec{\sigma}_{3} \right) Z_{12}(r_{12}) Y_{23}(r_{23}) - 2i Z_{12}(r_{12}) Z_{23}^{'}(r_{23}) \left(\vec{\sigma}_{2} + \vec{\sigma}_{3} \right) \cdot \hat{r}_{23} \times \nabla_{r_{2}^{'}} \right. \\ &+ 2i Z_{12}^{'}(r_{12}) Z_{23}^{'}(r_{23}) \left(\vec{\sigma}_{2} + \vec{\sigma}_{3} \right) \cdot \hat{r}_{12} \times \hat{r}_{23} + 2i Z_{12}(r_{12}) Z_{23}^{'}(r_{23}) \vec{\sigma}_{2} \cdot \hat{r}_{23} \times \nabla_{r_{3}^{'}} \\ &+ 2 Z_{12}^{'}(r_{12}) Z_{23}(r_{23}) \hat{r}_{12} \cdot \nabla_{r_{3}^{'}} + 2 Z_{12}(r_{12}) Z_{23}^{'}(r_{23}) \hat{r}_{23} \cdot \nabla_{r_{3}^{'}} + 4 Z_{12}(r_{12}) Z_{23}(r_{23}) \nabla_{r_{2}^{'}} \cdot \nabla_{r_{3}^{'}} \\ &- \frac{1}{3} \left(\vec{\sigma}_{2} \cdot \vec{\sigma}_{3} Y_{12}(r_{12}) + \hat{S}_{23}(\hat{r}_{23}) T_{23}(r_{23}) \right) Z_{12}(r_{12}) \right) \\ &+ D_{NNY}^{\sigma\omega,\bar{B}} \left(-Y_{12}(r_{12}) + Y_{23}(r_{23}) - 4 Z_{12}^{'}(r_{12}) Z_{23}^{'}(r_{23}) - 3 Z_{12}^{'}(r_{12}) \nabla_{r_{2}^{'}} \cdot \hat{r}_{12} \right) \\ &+ i \vec{\sigma}_{2} \cdot \left(2 \nabla_{r_{23}} \times \nabla_{r_{12}} - 5 \nabla_{r_{2}^{'}} \times \hat{r}_{23} \right) Z_{12}(r_{12}) Z_{23}(r_{23}) \\ &+ \left(\vec{r}_{12} \leftrightarrow \vec{r}_{23}, \vec{r}_{1}^{'} \leftrightarrow \vec{r}_{1}^{'}, \vec{r}_{2}^{'} \leftrightarrow \vec{r}_{2}^{'}, \vec{\sigma}_{1}^{'} \leftrightarrow \vec{\sigma}_{3} \right) \right) \delta(\vec{r}_{1}^{'} - \vec{r}_{1}^{'}) \delta(\vec{r}_{2}^{'} - \vec{r}_{2}^{'}) \delta(\vec{r}_{3}^{'} - \vec{r}_{3}^{'}) \end{split}$$

But that's only the beginning of the full story there are MANY, MANY, MANY more forces & contributions



Domenico in 2010

Domenico in 2013

BHF approximation of Hyperonic Matter

\diamond Energy per particle

•
$$\frac{E}{A}(\rho,\beta) = \frac{1}{A} \sum_{B} \sum_{k \le k_{F_B}} \left(\frac{\hbar^2 k^2}{2m_B} + \frac{1}{2} \operatorname{Re}\left[U_B(\vec{k}) \right] \right)$$

$$\infty$$

 \diamond Bethe-Goldstone Equation

•
$$G(\omega) = V + V \frac{Q}{\omega - E - E' + i\eta} G(\omega)$$

$$\bullet \quad E_B(k) = \frac{\hbar^2 k^2}{2m_B} + \operatorname{Re}\left[U_N(k)\right] + m_B$$

Partial sumation of pp ladder diagrams

$$\sum_{i=1}^{i} \sum_{j=1}^{i} \left(+ \right) = \left(+ \right) + \cdots =$$

$$= \sum_{i=1}^{i} \left(+ \right) + \left(+ \right) + \cdots =$$

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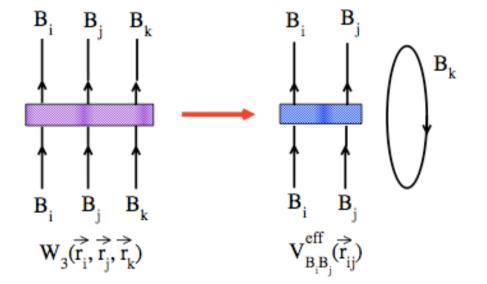
$$= \sum_{i=1}^{i} \left(+ \right) + \left(+ \right) + \cdots =$$

\diamond Coupled Channels

	S = 0 $S = -1$	S = -2	S = -3 S = -4
I = 0	$(NN \rightarrow NN)$	$ \begin{pmatrix} \Lambda\Lambda \to \Lambda\Lambda & \Lambda\Lambda \to \Xi N & \Lambda\Lambda \to \Sigma\Sigma \\ \Xi N \to \Lambda\Lambda & \Xi N \to \Xi N & \Xi N \to \Sigma\Sigma \\ \Sigma\Sigma \to \Lambda\Lambda & \Sigma\Sigma \to \Xi N & \Sigma\Sigma \to \Sigma\Sigma \end{cases} $	
I = 1/2	$\begin{pmatrix} \Lambda N \to \Lambda N & \Lambda \\ \Sigma N \to \Lambda N & \Sigma \end{pmatrix}$	$ \begin{array}{l} N \to \Sigma N \\ N \to \Sigma N \end{array} $	$\begin{pmatrix} \Lambda\Xi \to \Lambda\Xi & \Lambda\Xi \to \Sigma\Xi \\ \Sigma\Xi \to \Lambda\Xi & \Sigma\Xi \to \Sigma\Xi \end{pmatrix}$
I = 1	$(NN \rightarrow NN)$	$ \begin{pmatrix} \Xi N \to \Xi N & \Xi N \to \Lambda \Sigma & \Xi N \to \Sigma \Sigma \\ \Lambda \Sigma \to \Xi N & \Lambda \Sigma \to \Lambda \Sigma & \Lambda \Sigma \to \Sigma \Sigma \\ \Sigma \Sigma \to \Xi N & \Sigma \Sigma \to \Lambda \Sigma & \Sigma \Sigma \to \Sigma \Sigma \end{pmatrix} $	
I = 3/2	$(\Sigma N \rightarrow \Sigma$	N)	$(\Sigma\Xi \rightarrow \Sigma\Xi)$
I = 2		$(\Sigma\Sigma \rightarrow \Sigma\Sigma)$	

Three-Body Forces within the BHF approach

TBF can be introduced in our BHF approach by adding effective density-dependent two body forces to the baryon-baryon interactions V when solving the Bethe-Goldstone equation



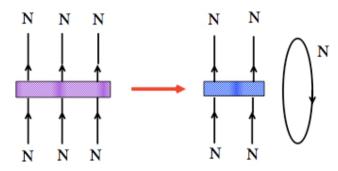
$$V_{B_iB_j}^{eff}\left(\vec{r}_{ij}\right) = \int W_3\left(\vec{r}_i,\vec{r}_j,\vec{r}_k\right) n\left(\vec{r}_i,\vec{r}_j,\vec{r}_k\right) d^3\vec{r}_k$$

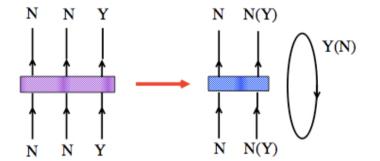
 $W_3(\vec{r}_i, \vec{r}_j, \vec{r}_k)$: genuine TBF $n(\vec{r}_i, \vec{r}_j, \vec{r}_k)$: three-body correlation function

From the genuine NNN,NNY, NYY and YYY TBF ...

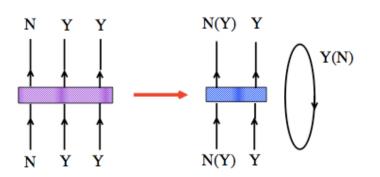




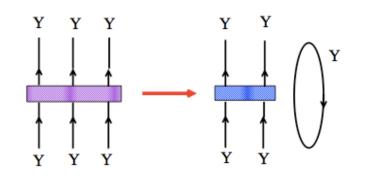


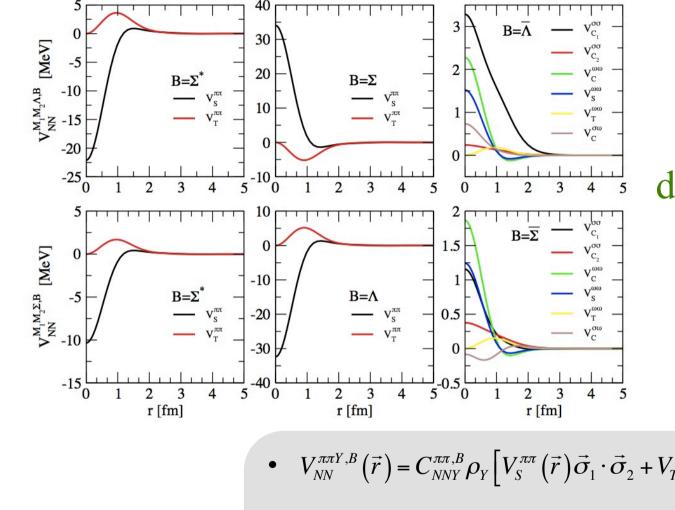


NYY → NY, YY



 $YYY \rightarrow YY$

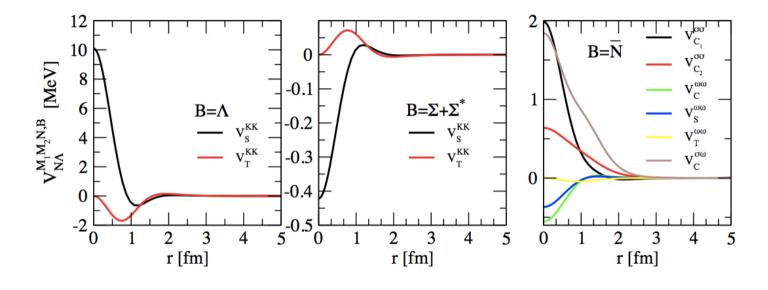




Effective NN density-dependent 2BF from NNY

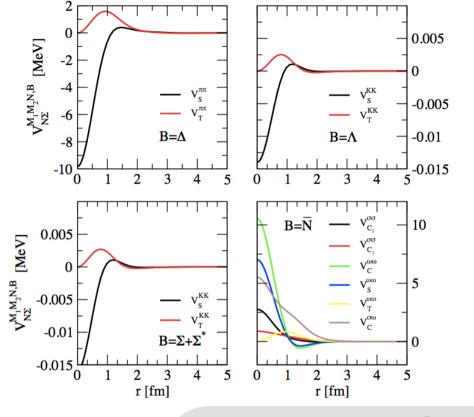
- $V_{NN}^{\pi\pi Y,B}\left(\vec{r}\right) = C_{NNY}^{\pi\pi,B}\rho_{Y}\left[V_{S}^{\pi\pi}\left(\vec{r}\right)\vec{\sigma}_{1}\cdot\vec{\sigma}_{2} + V_{T}^{\pi\pi}\left(\vec{r}\right)S_{12}\left(\hat{r}\right)\right]\vec{\tau}_{1}\cdot\vec{\tau}_{2}$
- $V_{NN}^{\sigma\sigma Y,\bar{B}}(\vec{r}) = C_{NNY}^{\sigma\sigma,\bar{B}} \left[\rho_N V_{C_1}^{\sigma\sigma}(\vec{r}) + \rho_N^{5/3} V_{C_2}^{\sigma\sigma}(\vec{r}) \right]$
- $V_{NN}^{\omega\omega Y,\bar{B}}\left(\vec{r}\right) = C_{NNY}^{\omega\omega,\bar{B}}\rho_{Y}\left[V_{C}^{\omega\omega}\left(\vec{r}\right) + V_{S}^{\omega\omega}\left(\vec{r}\right)\vec{\sigma}_{1}\cdot\vec{\sigma}_{2} + V_{T}^{\omega\omega}\left(\vec{r}\right)S_{12}\left(\hat{r}\right)\right]$
- $V_{NN}^{\sigma\omega Y,\bar{B}}(\vec{r}) = C_{NNY}^{\sigma\omega,\bar{B}}\rho_N V_C^{\sigma\omega}(\vec{r})$

Effective NA density-dependent 2BF from NNA



- $V_{N\Lambda}^{KKN,\Lambda}(\vec{r}) = C_{NN\Lambda}^{KK,\Lambda}\rho_N \left[V_S^{KK}(\vec{r})\vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{KK}(\vec{r})S_{12}(\hat{r}) \right]$
- $V_{N\Lambda}^{KKN,\Sigma/\Sigma^*}(\vec{r}) = C_{NN\Lambda}^{KK,\Sigma/\Sigma^*}\rho_N \left[V_S^{KK}(\vec{r})\vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{KK}(\vec{r})S_{12}(\hat{r})\right]\vec{\tau}_1 \cdot \vec{l}_2$
- $V_{N\Lambda}^{\sigma\sigma N,\bar{N}}(\vec{r}) = C_{NN\Lambda}^{\sigma\sigma,\bar{N}} \left[\rho_{\Lambda} V_{C_1}^{\sigma\sigma}(\vec{r}) + \rho_{\Lambda}^{5/3} V_{C_2}^{\sigma\sigma}(\vec{r}) \right]$
- $V_{N\Lambda}^{\omega\omega N,\bar{N}}\left(\vec{r}\right) = C_{NN\Lambda}^{\omega\omega,\bar{N}}\rho_{N}\left[V_{C}^{\omega\omega}\left(\vec{r}\right) + V_{S}^{\omega\omega}\left(\vec{r}\right)\vec{\sigma}_{1}\cdot\vec{\sigma}_{2} + V_{T}^{\omega\omega}\left(\vec{r}\right)S_{12}\left(\hat{r}\right)\right]$

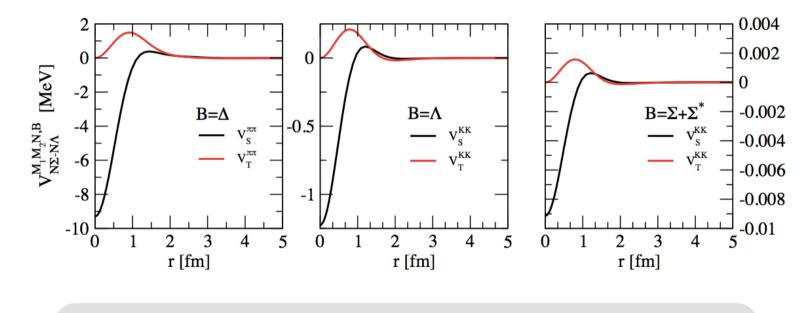
•
$$V_{N\Lambda}^{\sigma\omega N,\bar{N}}(\vec{r}) = C_{NN\Lambda}^{\sigma\omega,\bar{N}}\rho_{\Lambda}V_{C}^{\sigma\omega}(\vec{r})$$



Effective NΣ density-dependent 2BF from NNΣ

- $V_{N\Sigma}^{\pi\pi N,\Delta}(\vec{r}) = C_{NN\Sigma}^{\pi\pi,\Delta} \rho_N \left[V_S^{\pi\pi}(\vec{r}) \vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{\pi\pi}(\vec{r}) S_{12}(\hat{r}) \right] \vec{\tau}_1 \cdot \vec{I}_2$
- $V_{N\Sigma}^{KKN,\Lambda/\Sigma}(\vec{r}) = C_{NN\Sigma}^{KK,\Lambda/\Sigma}\rho_N \left[V_S^{KK}(\vec{r})\vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{KK}(\vec{r})S_{12}(\hat{r})\right]\vec{\tau}_1 \cdot \vec{\tau}_2$
- $V_{N\Sigma}^{KKN,\Sigma^*}(\vec{r}) = C_{NN\Sigma}^{KK,\Sigma^*}\rho_N \left[V_S^{KK}(\vec{r})\vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{KK}(\vec{r})S_{12}(\hat{r}) \right] \vec{\tau}_1 \cdot \vec{l}_2$
- $V_{N\Sigma}^{\sigma\sigma N,\bar{N}}\left(\vec{r}\right) = C_{NN\Sigma}^{\sigma\sigma,\bar{N}}\left[\rho_{\Sigma}V_{C_{1}}^{\sigma\sigma}\left(\vec{r}\right) + \rho_{\Sigma}^{5/3}V_{C_{2}}^{\sigma\sigma}\left(\vec{r}\right)\right]$
- $V_{N\Sigma}^{\omega\omega N,\bar{N}}(\vec{r}) = C_{NN\Sigma}^{\omega\omega,\bar{N}}\rho_N \left[V_C^{\omega\omega}(\vec{r}) + V_S^{\omega\omega}(\vec{r})\vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{\omega\omega}(\vec{r})S_{12}(\hat{r}) \right]$
- $V_{N\Sigma}^{\sigma\omega N,\bar{N}}\left(\vec{r}\right) = C_{NN\Sigma}^{\sigma\omega,\bar{N}}\rho_{\Sigma}V_{C}^{\sigma\omega}\left(\vec{r}\right)$

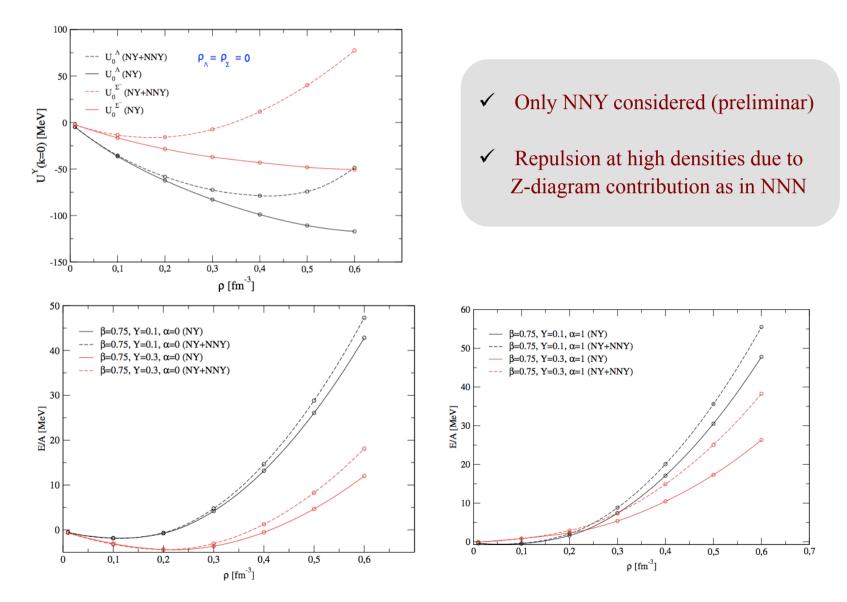
Effective density-dependent transition $N\Sigma - N\Lambda$ from $NN\Sigma - NN\Lambda$



• $V_{N\Sigma \leftrightarrow N\Lambda}^{\pi\pi N,\Delta}(\vec{r}) = C_{NN\Sigma \leftrightarrow NN\Lambda}^{\pi\pi,\Delta} \rho_N \left[V_S^{\pi\pi}(\vec{r}) \vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{\pi\pi}(\vec{r}) S_{12}(\hat{r}) \right] \vec{\tau}_1 \cdot \vec{I}_2$

•
$$V_{N\Sigma \leftrightarrow N\Lambda}^{KKN,\Lambda/\Sigma/\Sigma^*}(\vec{r}) = C_{NN\Sigma \leftrightarrow NN\Lambda}^{KK,\Lambda/\Sigma/\Sigma^*}\rho_N \left[V_S^{KK}(\vec{r})\vec{\sigma}_1 \cdot \vec{\sigma}_2 + V_T^{KK}(\vec{r})S_{12}(\hat{r})\right]\vec{\tau}_1 \cdot \vec{1}_2$$

Effect of TBF on Mean Field & E/A



Work is in progress, many more contributions have to be considered, but we can still try to estimate the effect of hyperonic TBF in NS

1-. Construct the hyperonic matter EoS within the BHF at 2 body level (Av18 NN + NSC89 YN)

2-. Add simple phenomenological density-dependent contact terms that mimic the effect of TBF.

Density-dependent contact terms: (Balberg & Gal 1997)

Potential of a baryon B_y in a sea of baryons B_x of density ρ_x

Folding $V_y(\rho_x)$ with ρ_x , $V_x(\rho_y)$ with ρ_y and combining with weight factors ρ_x / ρ and ρ_v / ρ

 $V_{y}(\rho_{x}) = a_{xy}\rho_{x} + b_{xy}\rho_{x}^{\gamma_{xy}}$

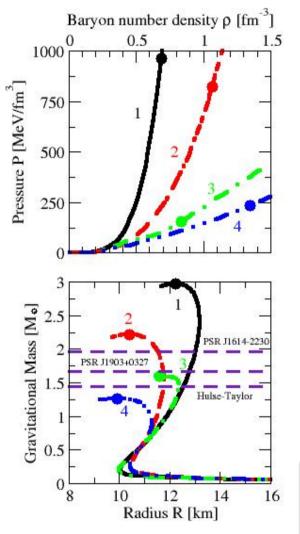
$$\varepsilon_{xy}(\rho_x,\rho_y) = a_{xy}\rho_x\rho_y + b_{xy}\rho_x\rho_y \left(\frac{\rho_x^{\gamma_{xy}} + \rho_y^{\gamma_{xy}}}{\rho_x + \rho_y}\right)$$

attraction



larger than 1

Effect of hyperonic TBF on M_{max}



$\frac{x}{2}$	γ_{YN}	Maximum Mass
0		
0	-	1.27(2.22)
1/3	1.49	1.33
2/3	1.69	1.38
1	1.77	1.41
0	-	1.29(2.46)
1/3	1.84	1.38
2/3	2.08	1.44
1	2.19	1.48
0	-	1.34(2.72)
1/3	2.23	1.45
2/3	2.49	1.50
1	2.62	1.54
0	-	1.38 (2.97)
1/3	2.63	1.51
2/3	2.91	1.56
1	3.05	1.60
	2/3 1 0 1/3 2/3 1 0 1/3 2/3 1 0 1/3	$\begin{array}{ccccccc} 2/3 & 1.69 \\ 1 & 1.77 \\ \hline 0 & - \\ 1/3 & 1.84 \\ 2/3 & 2.08 \\ 1 & 2.19 \\ \hline 0 & - \\ 1/3 & 2.23 \\ 2/3 & 2.49 \\ 1 & 2.62 \\ \hline 0 & - \\ 1/3 & 2.63 \\ 2/3 & 2.91 \\ \end{array}$

Hyperonic TBFs seem not to be the full solution of the "Hyperon Puzzle", although they probably contribute to its solution

 $1.27 < M_{\rm max} < 1.6 M_{\odot}$

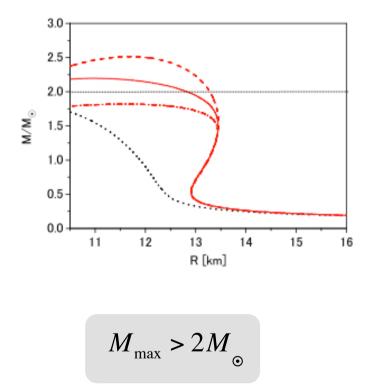


A comment must be done at this point



Yamamoto et al. (2015)

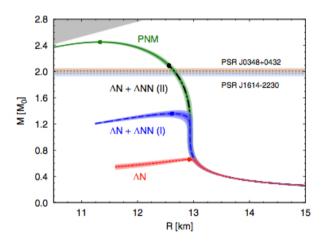
BHF with NN+YN+universal repulsive TBF (multipomeron exchange mecanism)





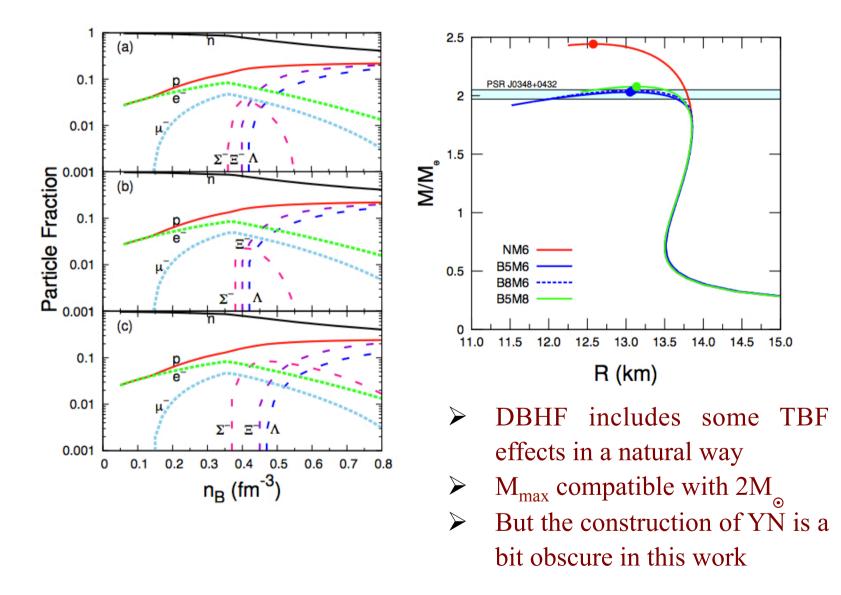
Lonardoni et al. (2015)

First Quantum Monte Carlo calculation on neutron+ Λ matter



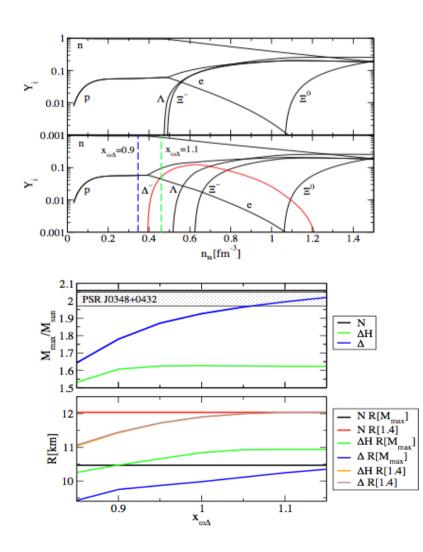
Some of the parametrizations of the Ann force give maximum masses compatible with 2M but the onset of A is above the largest density considered in the calculation (~ 0.56 fm⁻³). So in fact no As are present in NS interior

It should be mentioned also the recent DBHF calculation of hyperonic matter by Katayama & Saito (2014)



Is there also a Δ isobar puzzle ?

The recent work by Drago et al. (2014) calculation have studied the role of the Δ isobar in neutron star matter



- Constraints from L indicate an early appearance of Δ isobars in neutron stars matter at ~ 2-3 ρ_0 (same range as hyperons)
- Appearance of ∆ isobars modify the composition & structure of hadronic stars
- M_{max} is dramatically affected by the presence of Δ isobars

If Δ potential is close to that indicated by π -, e-nucleus or photoabsortion nuclear reactions then EoS is too soft $\longrightarrow \Delta$ puzzle similar to the hyperon one

Summary & Conclusions

Construction of two-meson exchange hyperonic TBF

Repulsion is obtained at high densities (Z-diagram)

D. Logoteta, Ph.D. Thesis (Univ. Coimbra, Sept. 2013)

Simple model to establish numerical lower and upper limits to the effect of hyperonicTBF on the maximum mass of NS.

Assuming the strength of hyperonic TBF \leq nucleonic TBF:

 $1.27 \text{ M}_{\odot} < M_{\text{max}} < 1.60 \text{ M}_{\odot}$ compatible with 1.4-1.5 M_{\odot}

but incompatible with observation of very massive NS

 $\begin{array}{l} \text{PSR J1903+0327} \quad (1.67 \pm 0.01) \ \text{M}_{\odot} \\ \text{PSR J1614-2230} \quad (1.97 \pm 0.04) \ \text{M}_{\odot} \\ \text{PSR J0348+0432} \quad (2.01 \pm 0.04) \ \text{M}_{\odot} \end{array}$

There is not yet a general agreement between different approaches/models

Take away message



Hyperonic Three-Body Forces seem not to be the full solution to the "Hyperon Puzzle", although they probably can contribute to it

- You for your time & attention
- The organizers for their invitation & support



Nordic Institute for Theoretical Physics







